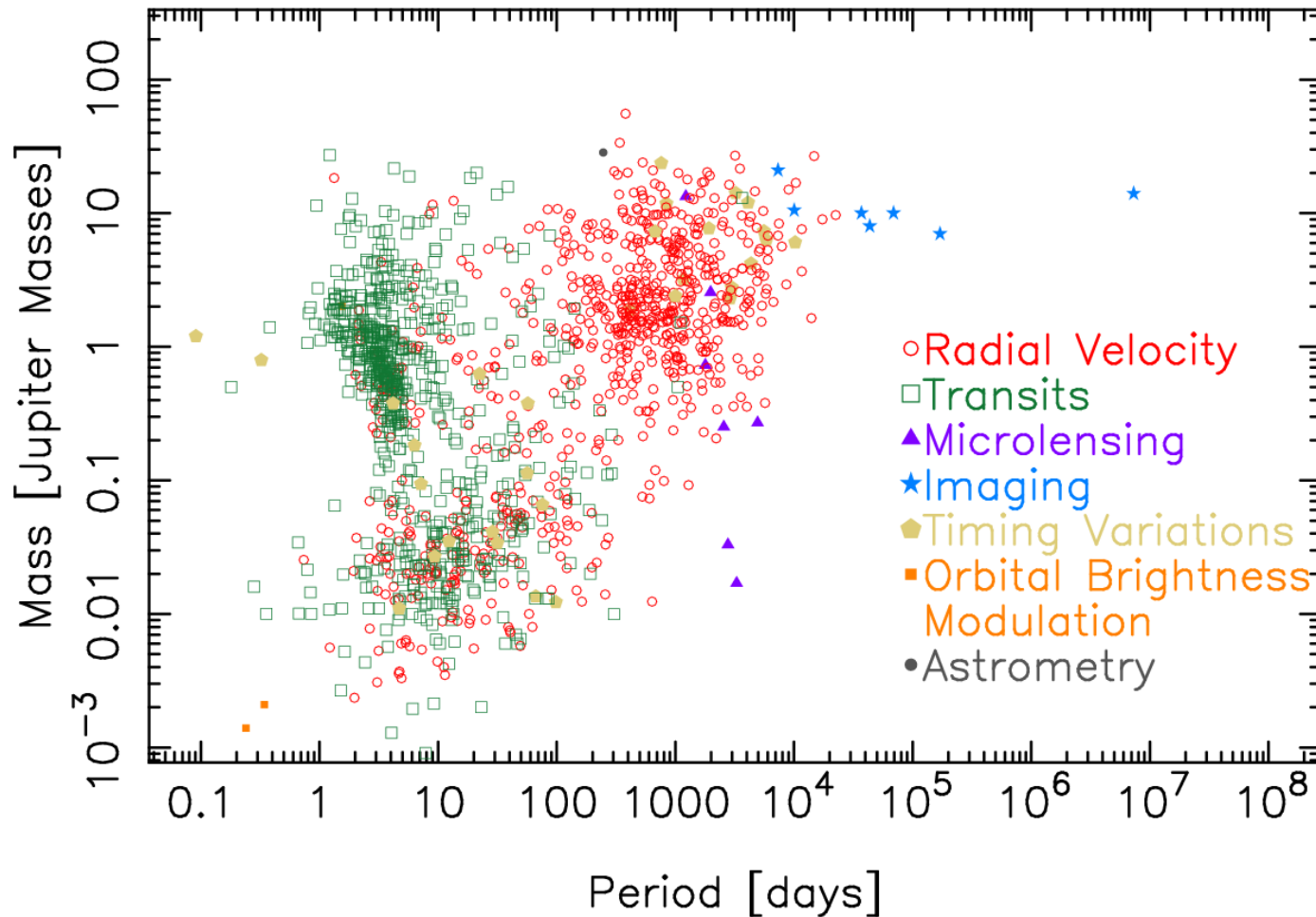


# TESS Observations of Known Exoplanetary Systems

25 Jul 2019  
exoplanetarchive.ipac.caltech.edu

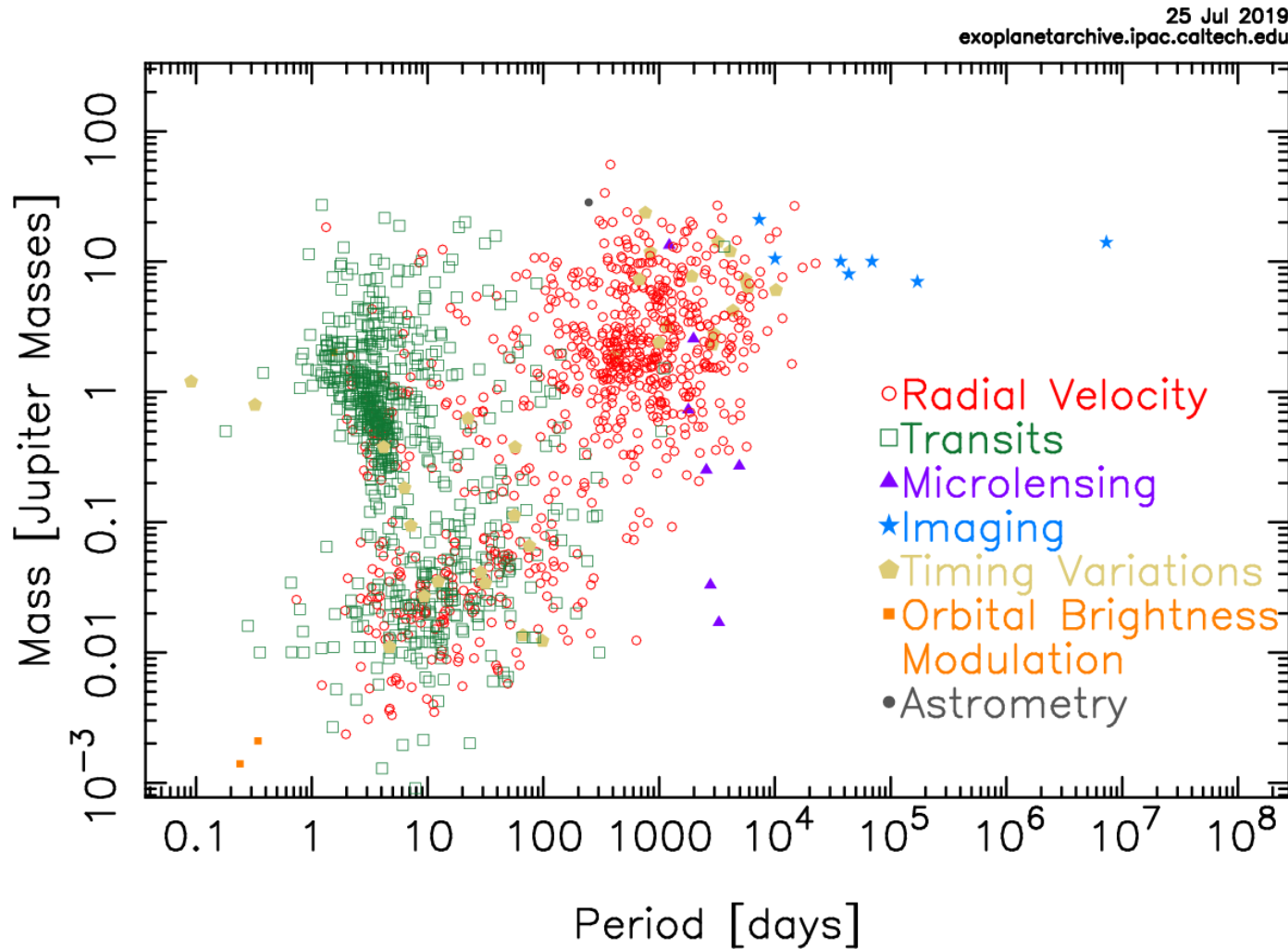


**Stephen Kane, Jacob Bean, Tiago Campante, Paul Dalba,  
Tara Fetherolf, Greg Henry, Teo Mocnik, Joshua Pepper,  
Darin Ragozzine, Maggie Turnbull**

Total number of confirmed planets (as of July 25, 2019):

**4,025**

Discovery method: Transits 3,114; RVs 760

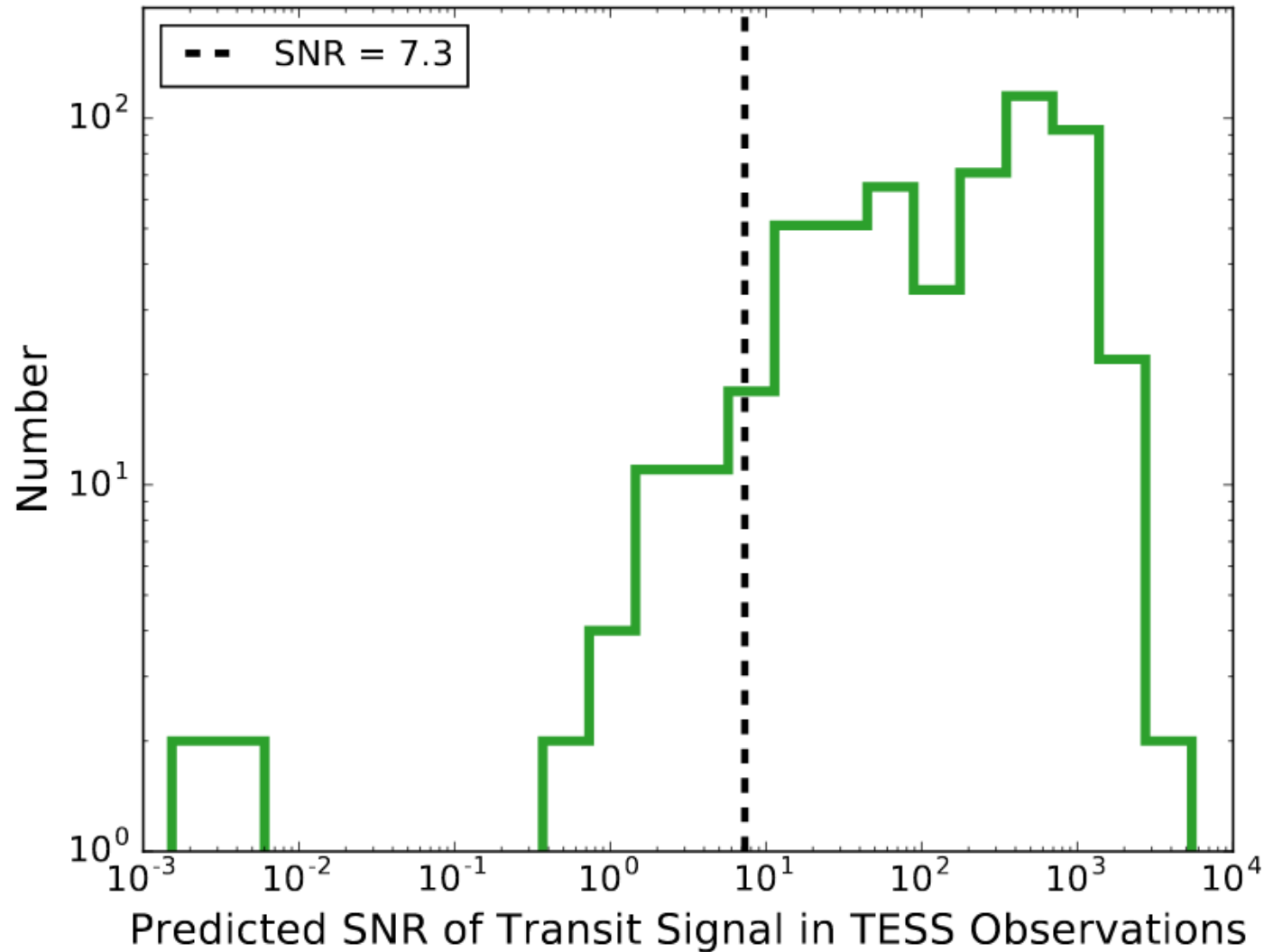


**During the TESS primary mission, 83% of known exoplanet hosts will be observed.  
(Using the TESS Viewing Tool by Tom Barclay)**

# Why Study the Known Exoplanets with TESS?

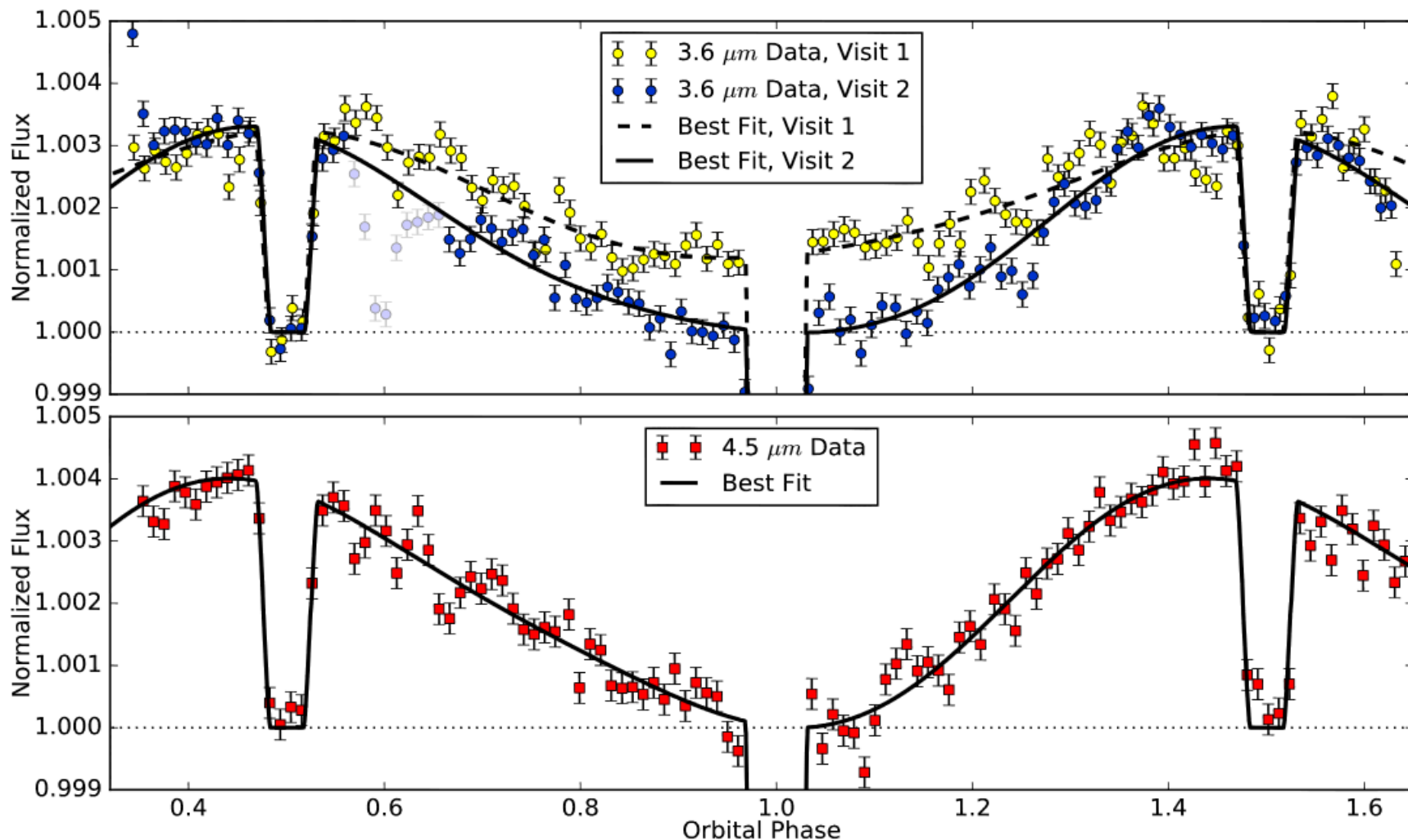
# Why Study the Known Exoplanets with TESS?

- Transit detection of known radial velocity planets.



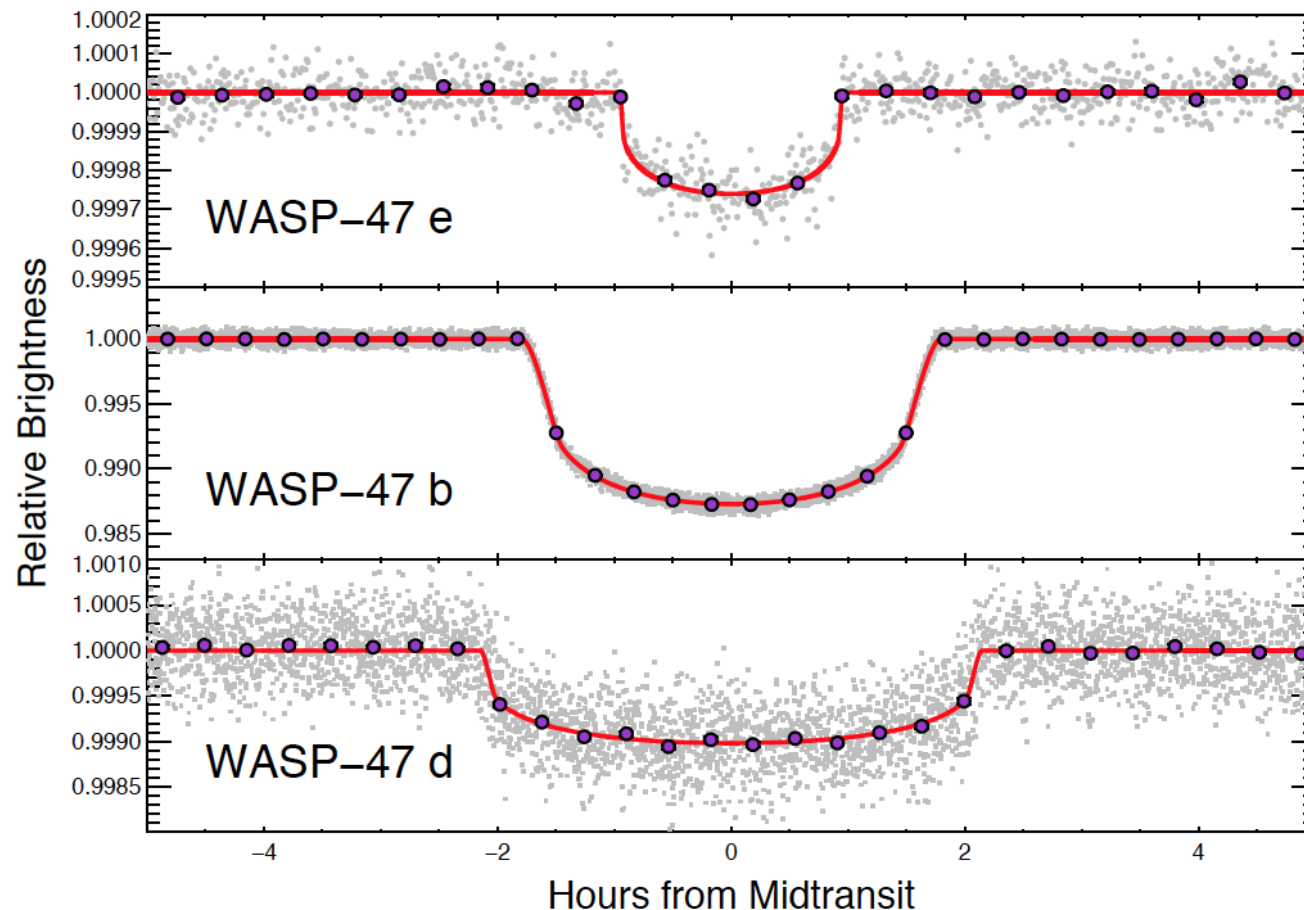
# Why Study the Known Exoplanets with TESS?

- Transit detection of known radial velocity planets.
- Detection of phase variations, and secondary eclipses for known planets.



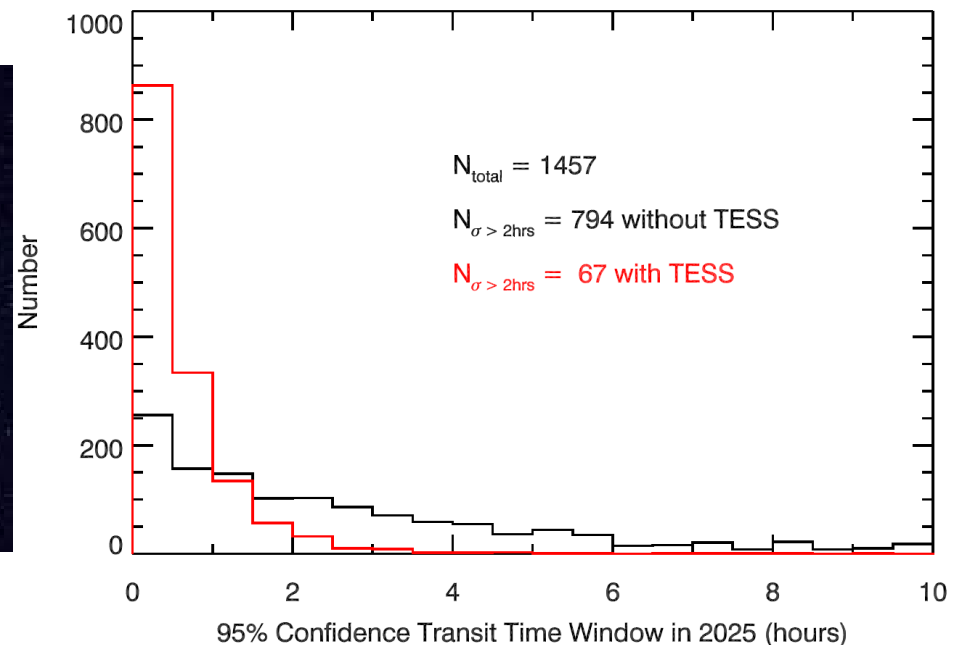
# Why Study the Known Exoplanets with TESS?

- Transit detection of known radial velocity planets.
- Detection of phase variations, and secondary eclipses for known planets.
- Transits of additional planets in known systems.



# Why Study the Known Exoplanets with TESS?

- Transit detection of known radial velocity planets.
- Detection of phase variations, and secondary eclipses for known planets.
- Transits of additional planets in known systems.
- Ephemeris refinement for known transitters:
  - Preparation for JWST observations.
  - Additional TTVs of known transiting planets.
  - Orbital decay for Hot Jupiters



# Why Study the Known Exoplanets with TESS?

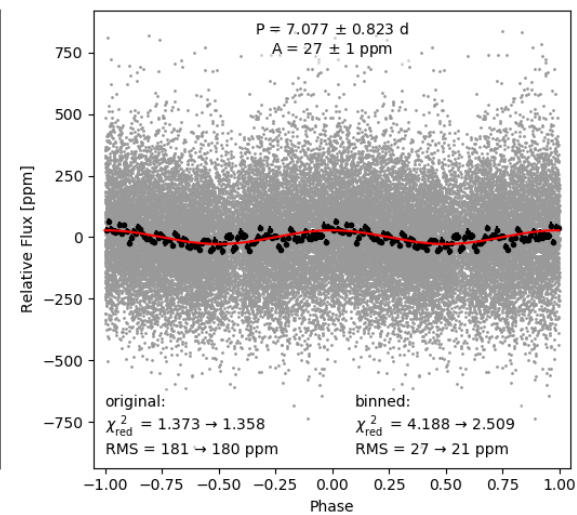
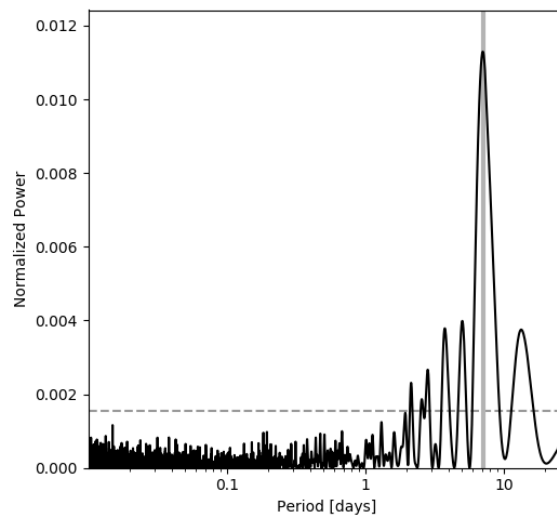
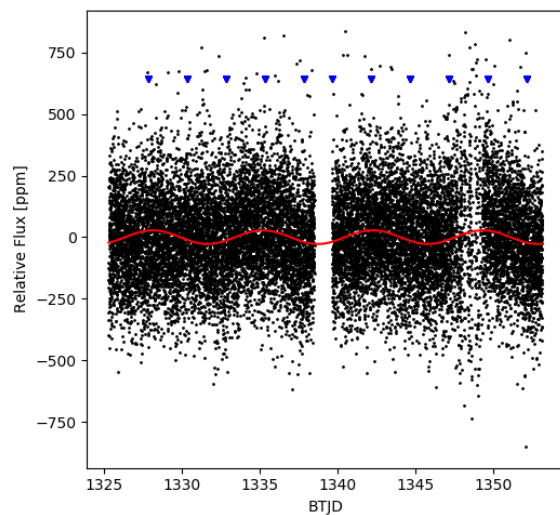
- Transit detection of known radial velocity planets.
- Detection of phase variations, and secondary eclipses for known planets.
- Transits of additional planets in known systems.
- Ephemeris refinement for known transitters:
  - Preparation for JWST observations.
  - Additional TTVs of known transiting planets.
  - Orbital decay for Hot Jupiters
- Astroseismology of known planet hosts.



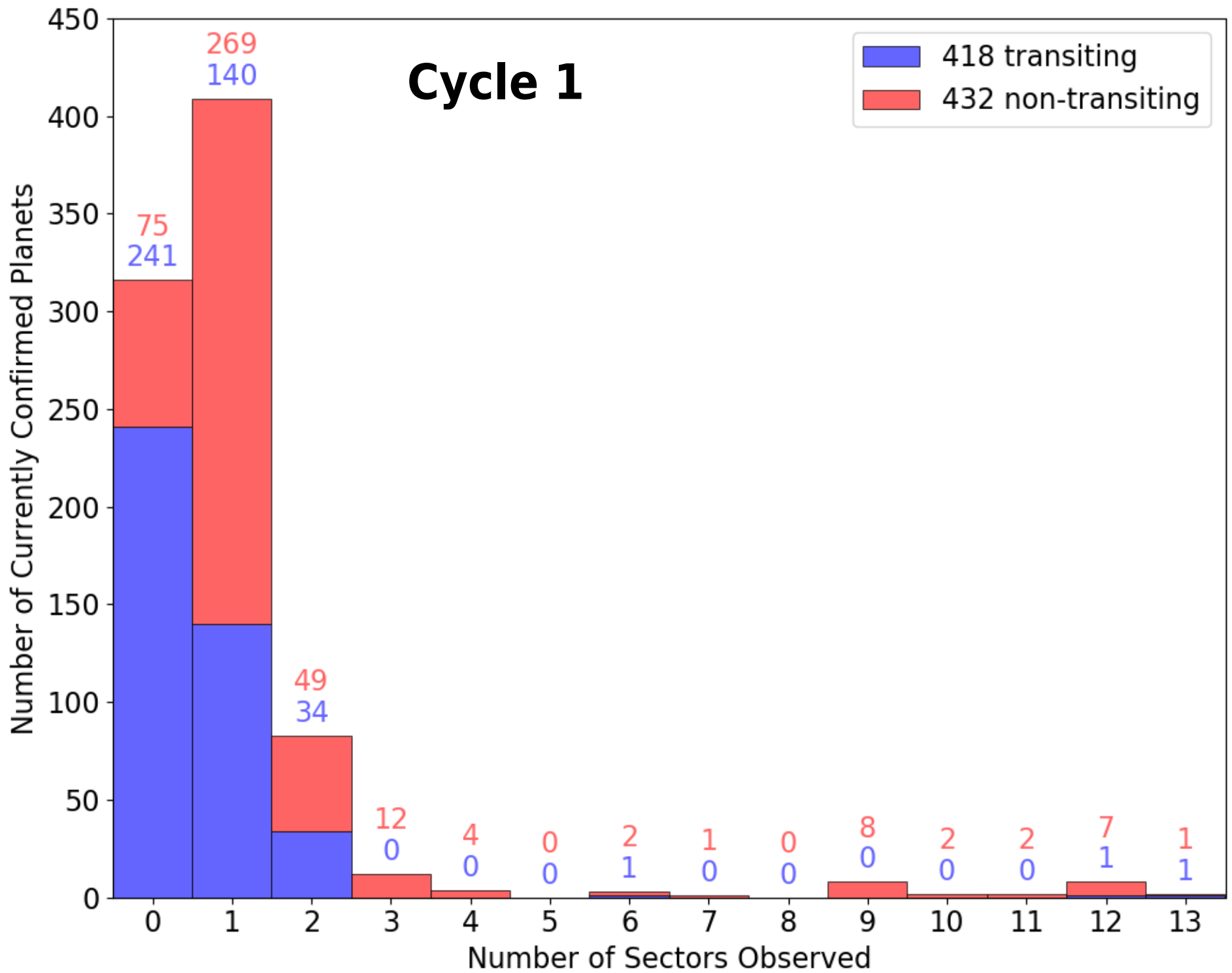


# Why Study the Known Exoplanets with TESS?

- Transit detection of known radial velocity planets.
- Detection of phase variations, and secondary eclipses for known planets.
- Transits of additional planets in known systems.
- Ephemeris refinement for known transitters:
  - Preparation for JWST observations.
  - Additional TTVs of known transiting planets.
  - Orbital decay for Hot Jupiters
- Astroseismology of known planet hosts.
- Intrinsic stellar variability of known exoplanet hosts.



# Cycle 1

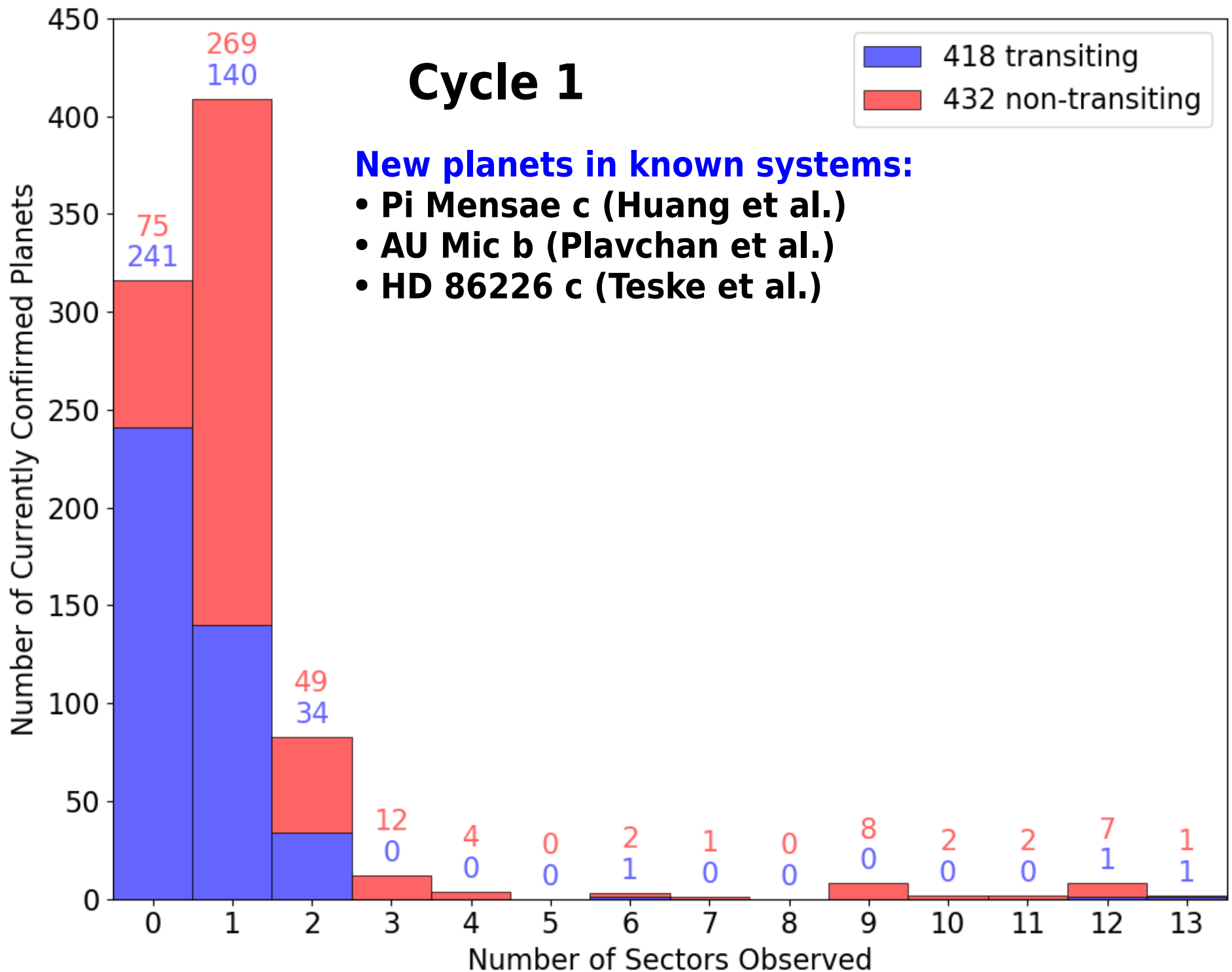


# Cycle 1

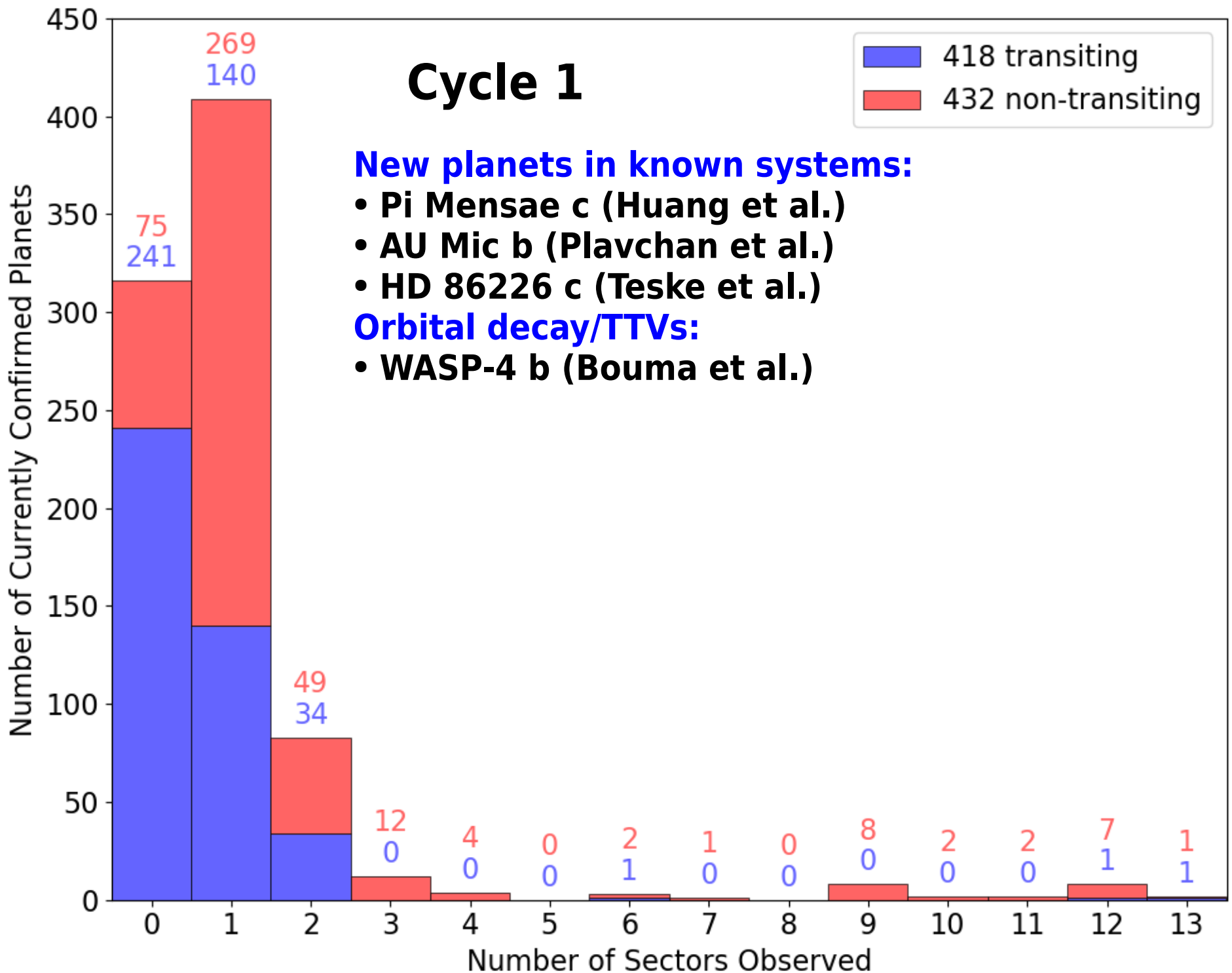
418 transiting  
432 non-transiting

## New planets in known systems:

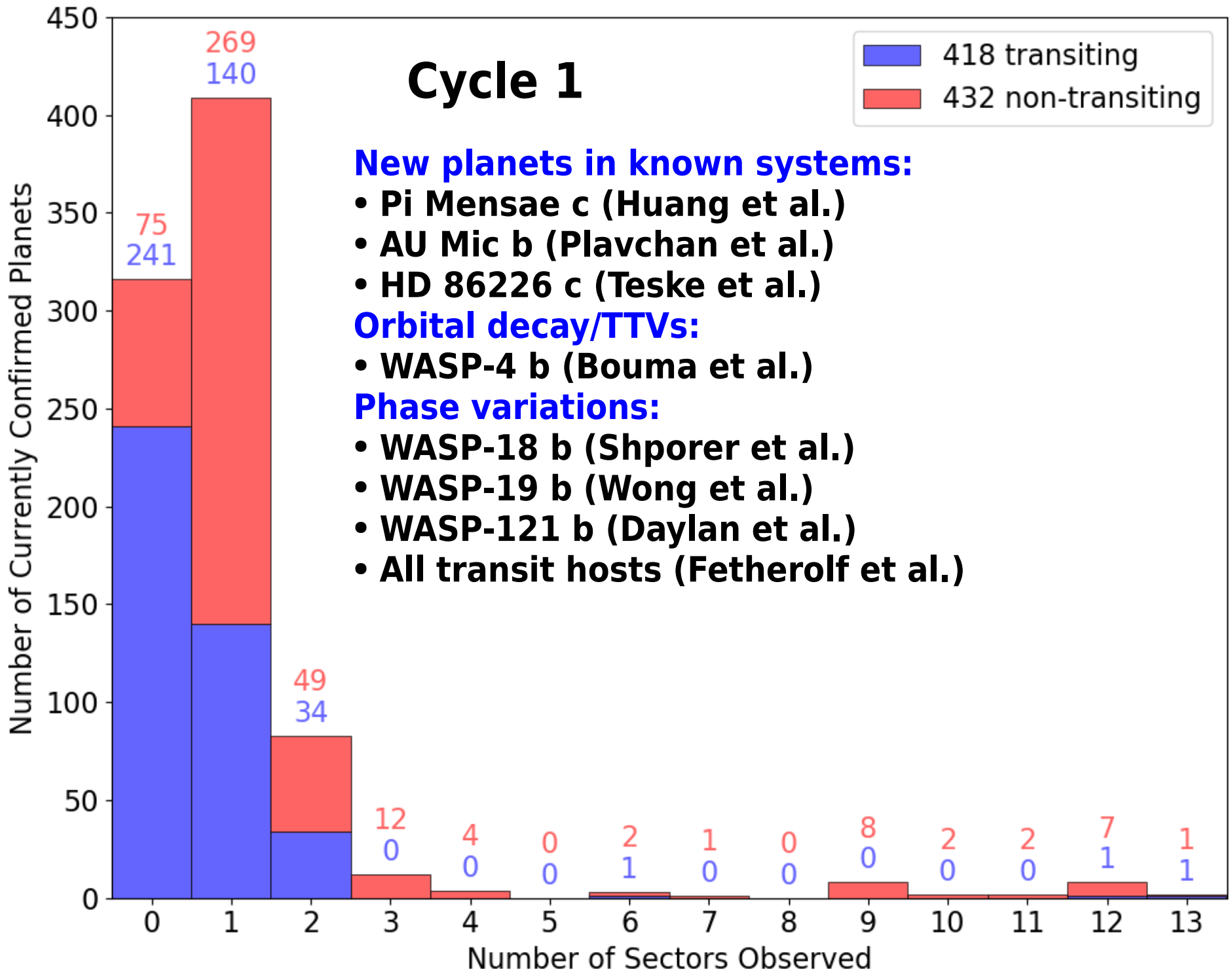
- Pi Mensae c (Huang et al.)
- AU Mic b (Plavchan et al.)
- HD 86226 c (Teske et al.)



# Cycle 1



# Cycle 1



## New planets in known systems:

- Pi Mensae c (Huang et al.)
- AU Mic b (Plavchan et al.)
- HD 86226 c (Teske et al.)

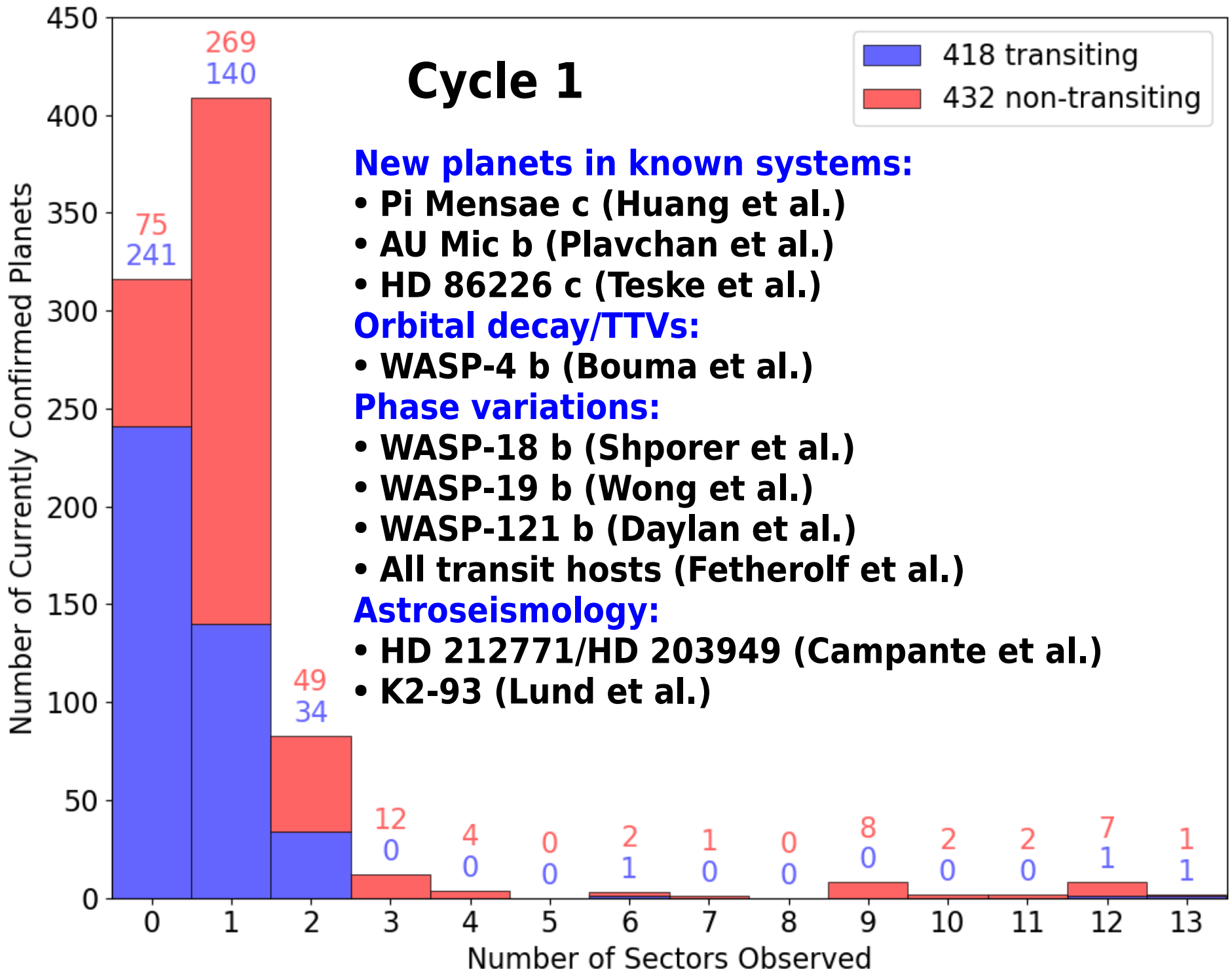
## Orbital decay/TTVs:

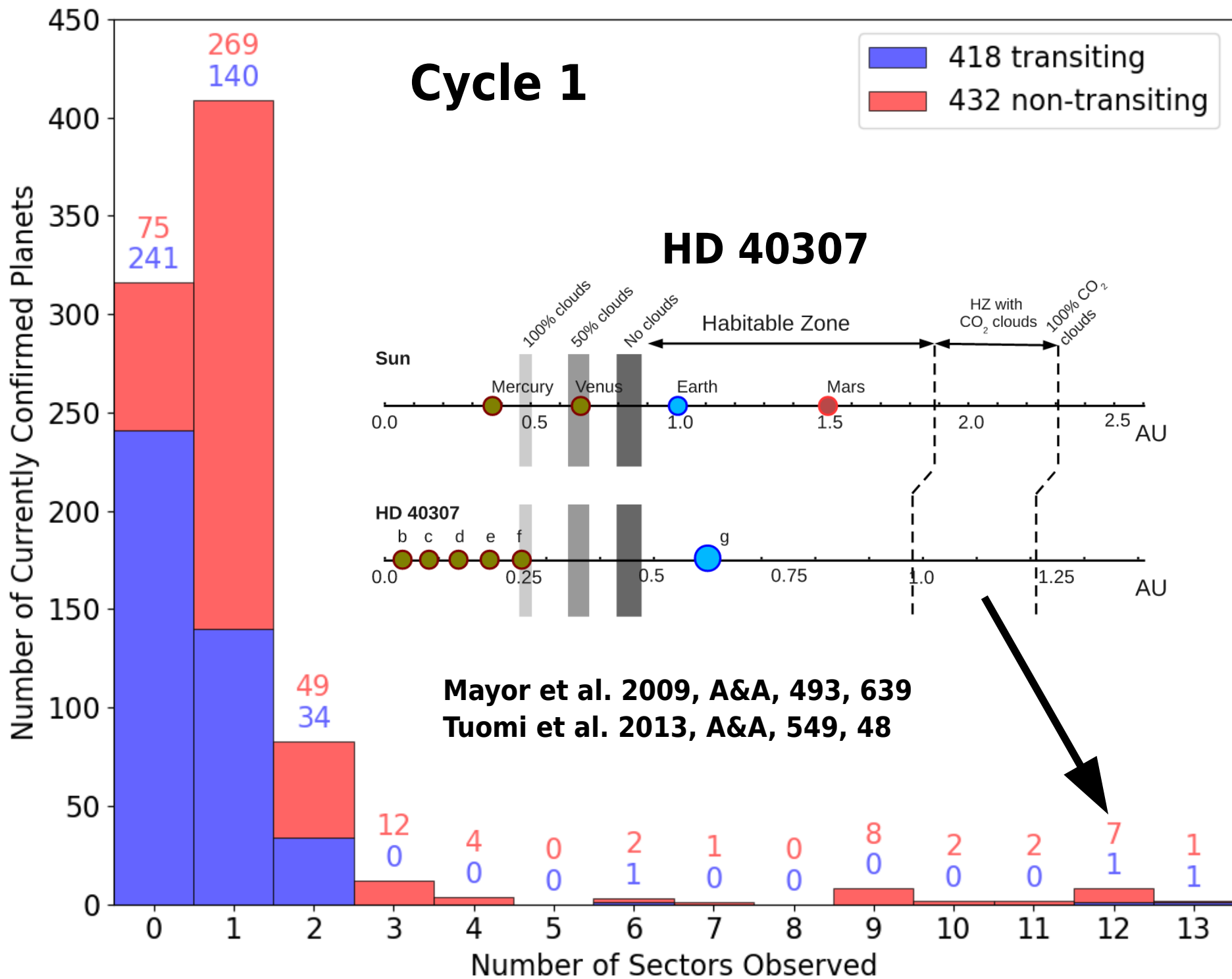
- WASP-4 b (Bouma et al.)

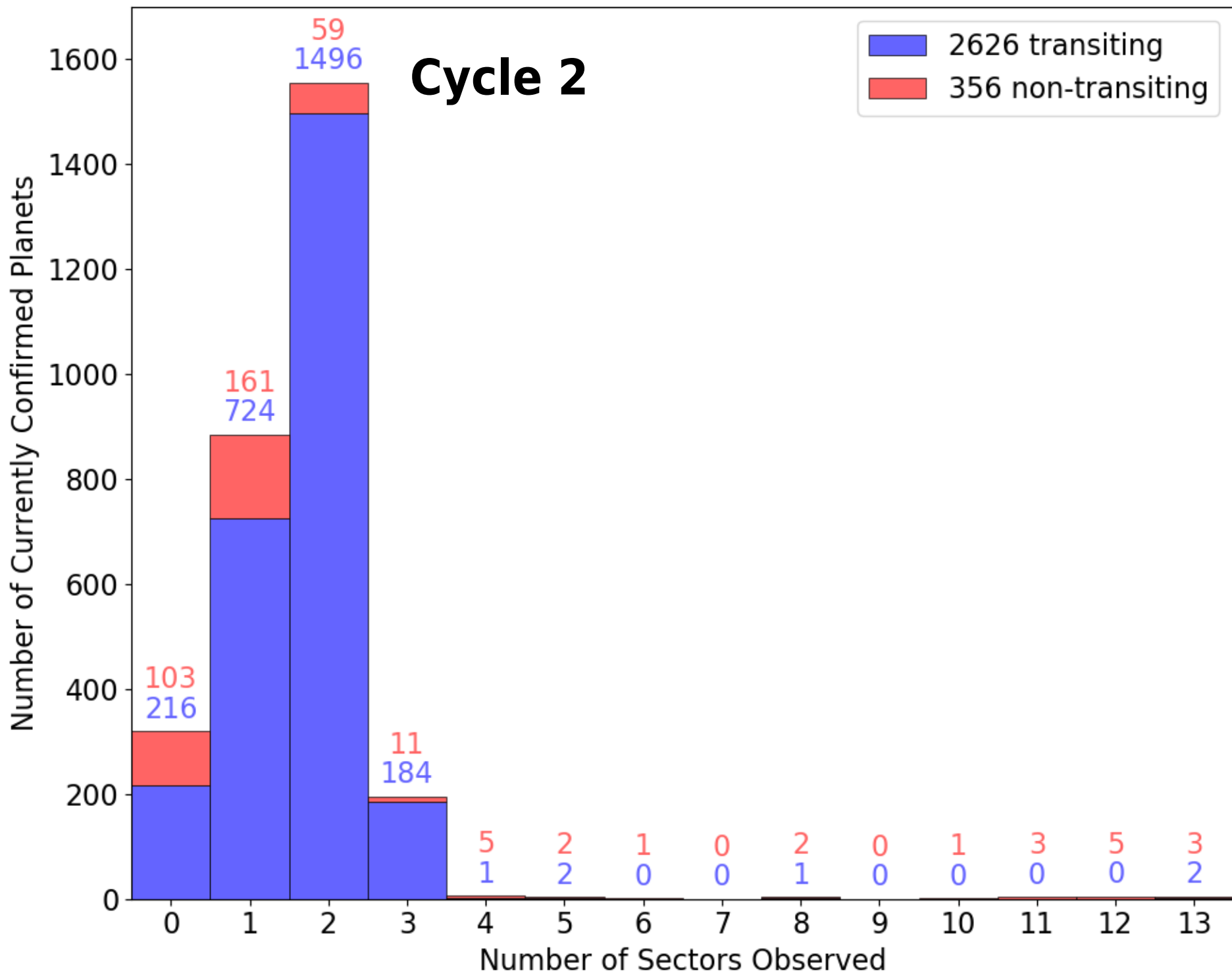
## Phase variations:

- WASP-18 b (Shporer et al.)
- WASP-19 b (Wong et al.)
- WASP-121 b (Daylan et al.)
- All transit hosts (Fetherolf et al.)

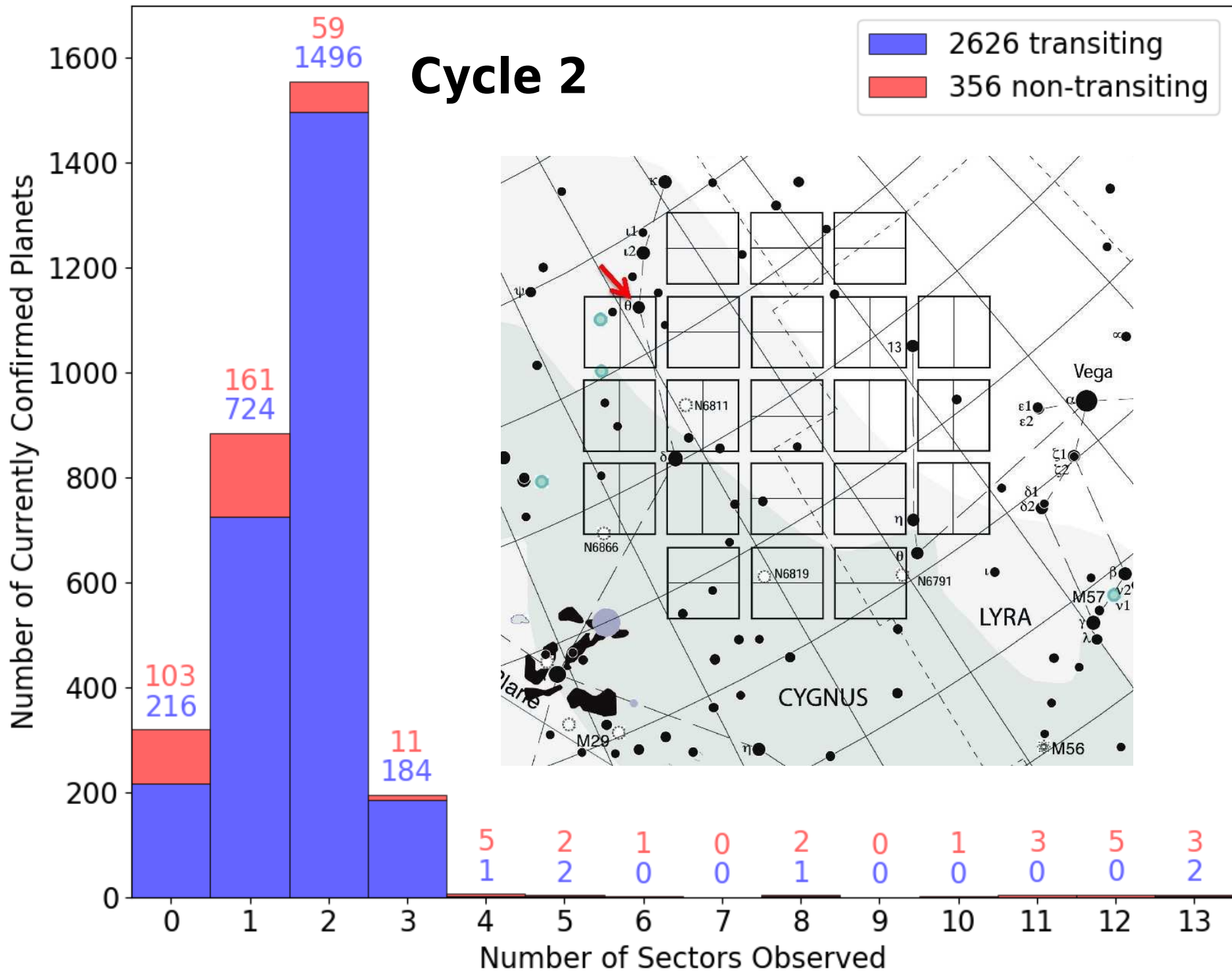
# Cycle 1











# Why Study the Known Exoplanets with TESS?

- **Transit detection of known radial velocity planets.**
- **Detection of phase variations, and secondary eclipses for known planets.**
- **Transits of additional planets in known systems.**
- **Ephemeris refinement for known transitters:**
  - **Preparation for JWST observations.**
  - **Additional TTVs of known transiting planets.**
  - **Orbital decay for Hot Jupiters**
- **Astroseismology of known planet hosts.**
- **Intrinsic stellar variability of known exoplanet hosts.**

**Feel free to ask me for known exoplanet data.**